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On a class of Ricci-Quadratic Finsler metrics

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Abstract. Let F be a (reversible) Finsler metric on a Riemannian space (M, α) of positive (or negative) sectional curvature. Suppose that the Ricci curvature of F is horizontally constant along Finslerian geodesics. Then we show that F is a Ricci-quadratic Finsler metric.

Keywords: Finsler space, Ricci-quadratic Finsler metric, **E**-curvature, **H**-curvature, Anisotropic space-time.

1. INTRODUCTION

An (α, β) -metric F is a Finsler metric on the background Riemannian manifold (M, α) . Therefore, on is dealing with two metrics F and α within the related computations. This bi-metric issue may be crucial for applied disciplines and there may be considered several types of bi-metric spaces. For example, the anisotropy property can be detected using radiation in the background Riemannian space. One may assume that the background Riemannian space has some specific geometric properties; Bi-metric theories in General Relativity are of such various types and contain both the usual metric and a metric of constant curvature, and may contain other scalar or vector fields, cf. [4].

Given a Finsler metric F = F(x, y), the locally minimizing curves are characterized by the system of differential equations

$$\ddot{c}^i(t) + 2G^i\big(c(t), \dot{c}(t)\big) = 0,$$

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where the local functions $G^i = G^i(x, y)$ are called the spray coefficients. For a Riemannian metric $F = \sqrt{g_{ij}(x)y^iy^j}$, the spray coefficients are quadratic in $y \in T_x M$. There are non-Riemannian metrics whose spray coefficients still have this quadratic property. Finsler metrics with this property are called Berwald metrics. In this case, we have

$$G^i = \frac{1}{2} \Gamma^i{}_{jk}(x) y^j y^k.$$

The Chern connection (as well as the Berwald connection) of any Berwald metric F is the Levi-Civita connection of a Riemannian metric α and the Riemann and the Ricci curvatures of F are eventually those of the Riemannian metric α . Hence every Berwald space deals with a bi-metrics theory.

The notion of Riemann curvature for Riemann metrics can be extend to Finsler metrics. For $y \in T_x M_0$, the Riemann curvature $\mathbf{R}_y : T_x M \to T_x M$ is defined by $\mathbf{R}_y(u) = R^i_{\ k}(y)u^k \frac{\partial}{\partial x^i}$ where

$$R^{i}{}_{k}(y) := 2\frac{\partial G^{i}}{\partial x^{k}} - \frac{\partial^{2} G^{i}}{\partial x^{j} \partial y^{k}}y^{j} + 2G^{j}\frac{\partial^{2} G^{i}}{\partial y^{j} \partial y^{k}} - \frac{\partial G^{i}}{\partial y^{j}}\frac{\partial G^{j}}{\partial y^{k}}.$$
 (1.1)

The Ricci curvature on an n-manifold M is defined by

$$\mathbf{Ric} = \sum_{k=1}^{n} R^{k}_{\ k}(x, y).$$

By definition, the Ricci curvature is a positively homogeneous function of degree two in $y \in TM$. But it is not quadratic in $y \in T_xM$, in general. From Eq.(1.1), one can see that if F is a Berwald metric then the Ricci curvature is quadratic in $y \in T_xM$. Finsler metrics with such curvature property are called *Ricciquadratic metrics* [9]. The key idea for Finsler metrics with positive quadratic Ricci curvature is that thereby the Ricci curvature

$$\mathbf{Ric}(x,y) = h_{ij}(x)y^i y^j$$

defines a natural Riemannian metric on M given by $h = h_{ij}(x)dx^i dx^j$.

The Randers metrics are the most popular Finsler metrics appearing in many areas of Differential geometry and Physics and simply accessible by a Riemannian metrics $\alpha = \sqrt{a_{ij}(x)y^iy^j}$ and a 1-form $\beta = b_i(x)y^i$ on a manifold M. It has been in the center of researches devoted in unified field theory for long years after G. Randers applied it in [10]. In [9], Li-Shen characterize Ricci-quadratic Randers metrics.

Let us denote the Levi-Civita connection of α by $\tilde{\nabla}$ and denote the horizontal and vertical covariant derivations with respect to the horizontal vector $\frac{\tilde{\delta}}{\delta x^i}$ and the vertical vector $\frac{\partial}{\partial y^i}$ associated to $\tilde{\nabla}$ by " $_{|_i}$ " and ";i" respectively. Let

$$\mathbf{Ric}_{ij} := rac{1}{2}\mathbf{Ric}_{;i;j} = rac{1}{2}rac{\partial^2\mathbf{Ric}}{\partial y^i \partial y^j}$$

where **Ric** is the Ricci tensor of F and " $_{|_0} := {}_{|_s}y^s$ " is denote the horizontal covariant derivation on geodesics of Riemannian metric α .

In this paper we prove the following result:

Theorem 1.1. Let F be a (reversible) Finsler metric on a background Riemannian space (M, α) of positive (or negative) sectional curvature. Suppose that Ricci curvature satisfies following

 $Ric_{ij|0} = 0.$

Then F is Ricci-quadratic.

There are many Finsler metrics whose Riemann curvature in every direction is quadratic. A Finsler metric F is said to be R-quadratic if R_y is quadratic in $y \in T_x M$ at each point $x \in M$. Indeed a Finsler metric is R-quadratic if and only if the h-curvature of Berwald connection depends on position only in the sense of Bácsó-Matsumoto [3]. We have $R_k^i = R_j^i_{kl}(x, y)y^jy^l$. Therefore R_k^i is quadratic in $y \in T_x M$ if and only if $R_j^i_{kl}$ are functions of position alone. In this case, we have

$$R_k^i = R_j^i_{kl}(x)y^j y^l$$

It is remarkable that, the notion of R-quadratic Finsler metrics was introduced by Shen, which can be considered as a generalization of Berwald metrics and R-flat metrics [20]. He proved that every compact R-quadratic Finsler metric is a Landsberg metric. In [16], Najafi-Bidabad-Tayebi showed that every Rquadratic Finsler metric satisfies $\mathbf{H} = 0$.

A Finsler metric F is said to be *Ricci-quadratic* if *Ricci* is quadratic in $y \in T_x M$ at each point $x \in M$. In this paper, we prove the following.

Theorem 1.2. Every Ricci-quadratic Finsler manifold (M, F) is of vanishing H-curvature.

2. Preliminaries

Let M be a n-dimensional C^{∞} manifold. Denote by $T_x M$ the tangent space at $x \in M$, by $TM = \bigcup_{x \in M} T_x M$ the tangent bundle of M and by $TM_0 := TM \setminus \{0\}$ the slit tangent bundle.

A Finsler metric on M is a function $F:TM\to [0,\infty)$ which has the following properties:

(i) F is C^{∞} on $TM_0 := TM \setminus \{0\};$

(ii) F is positively 1-homogeneous on the fibers of tangent bundle TM;

(iii) for each $y \in T_x M$, the following quadratic form $\mathbf{g}_y : T_x M \times T_x M \to \mathbb{R}$ on $T_x M$ is positive definite,

$$\mathbf{g}_{y}(u,v) := \frac{1}{2} \Big[F^{2}(y + su + tv) \Big]|_{s,t=0}, \quad u,v \in T_{x}M.$$

Let $x \in M$ and $F_x := F|_{T_xM}$. To measure the non-Euclidean feature of F_x , define $\mathbf{C}_y : T_xM \times T_xM \times T_xM \to \mathbb{R}$ by $\mathbf{C}_y(u, v, w) := C_{ijk}(y)u^iv^jw^k$ where

$$C_{ijk}(y) := \frac{1}{4} \frac{\partial^3 F^2}{\partial y^i \partial y^j \partial y^k}(y)$$

The family $\mathbf{C} := {\mathbf{C}_y}_{y \in TM_0}$ is called the *Cartan torsion*. It is well known that $\mathbf{C=0}$ if and only if F is Riemannian.

The horizontal covariant derivatives of **C** along geodesics give rise to the Landsberg curvature $\mathbf{L}_y: T_x M \times T_x M \times T_x M \to \mathbb{R}$ defined by

$$\mathbf{L}_{y}(u, v, w) := L_{ijk}(y)u^{i}v^{j}w^{k},$$

where $u = u^i \frac{\partial}{\partial x^i}|_x$, $v = v^i \frac{\partial}{\partial x^i}|_x$, $w = w^i \frac{\partial}{\partial x^i}|_x$ and $L_{ijk} := C_{ijk|s}y^s$. The family $\mathbf{L} := {\mathbf{L}_y}_{y \in TM_0}$ is called the *Landsberg curvature*. A Finsler metric is called a *Landsberg metric* if $\mathbf{L=0}$ [18].

Given a Finsler manifold (M, F), then a global vector field **G** is induced by F on TM_0 , which in a standard coordinate (x^i, y^i) for TM_0 is given by

$$\mathbf{G} = y^i \frac{\partial}{\partial x^i} - 2G^i \frac{\partial}{\partial y^i},$$

where $G^i = G^i(x, y)$ are local functions on TM given by

$$G^{i} := \frac{1}{4}g^{il} \left\{ \frac{\partial^{2}[F^{2}]}{\partial x^{k} \partial y^{l}} y^{k} - \frac{\partial [F^{2}]}{\partial x^{l}} \right\}, \quad y \in T_{x}M.$$

G is called the associated spray to (M, F).

For $y \in T_x M_0$, define $\mathbf{B}_y : T_x M \otimes T_x M \otimes T_x M \to T_x M$ and $\mathbf{E}_y : T_x M \otimes T_x M \to \mathbb{R}$ by

$$\mathbf{B}_{y}(u,v,w) := B^{i}{}_{jkl}(y)u^{j}v^{k}w^{l}\frac{\partial}{\partial x^{i}}|_{x}, \quad \mathbf{E}_{y}(u,v) := E_{jk}(y)u^{j}v^{k},$$

where

$$B^{i}_{\ jkl}(y) := \frac{\partial^{3} G^{i}}{\partial y^{j} \partial y^{k} \partial y^{l}}(y), \quad E_{jk}(y) := \frac{1}{2} B^{m}_{\ jkm}(y).$$

B and **E** are called the Berwald curvature and mean Berwald curvature respectively. A Finsler metric F is called a Berwald metric and weakly Berwald metric if $\mathbf{B} = 0$ and $\mathbf{E} = 0$, respectively [19].

The quantity $\mathbf{H}_y = H_{ij} dx^i \otimes dx^j$ is defined as the covariant derivative of \mathbf{E} along geodesics. More precisely

 $H_{ij} := E_{ij|m} y^m.$

For H_{ij} , we get $H_{ij}y^i = 0$ (see [1], [17] and [25]).

Azadeh Shirafkan

The Busemann-Hausdorff volume form $dV_F = \sigma_F(x)dx^1 \cdots dx^n$ on any Finsler space (M, F) is defined by

$$\sigma_F(x) := \frac{\operatorname{Vol}(\mathbb{B}^n(1))}{\operatorname{Vol}\left\{(y^i) \in \mathbb{R}^n \mid F(y^i \frac{\partial}{\partial x^i}|x) < 1\right\}}.$$

Assume that

$$\underline{g} = \det\left(g_{ij}(x,y)\right)$$

and define

$$\tau(x,y) := \ln \frac{\sqrt{g}}{\sigma_F(x)}.$$

Then, $\tau = \tau(x, y)$ is a scalar function on slit tangent bundle TM_0 , which is called the *distortion* [19].

For a vector $\mathbf{y} \in T_x M$, let $c(t), -\epsilon < t < \epsilon$, denote the geodesic with c(0) = xand $\dot{c}(0) = \mathbf{y}$. The function

$$\mathbf{S}(\mathbf{y}) := \frac{d}{dt} \Big[\tau(\dot{c}(t)) \Big]|_{t=0}$$

is called the S-curvature with respect to the Busemann-Hausdorff volume form.

A Finsler space is said to be *of isotropic* **S***-curvature* if there is a function c = c(x) defined on M such that

$$\mathbf{S} = (n+1)c(x)F.$$

It is called a Finsler space of constant \mathbf{S} -curvature once c is a constant. Every Berwald space is of vanishing \mathbf{S} -curvature [19]. Notice that, \mathbf{S} -curvature are in fact non-Riemannian quantities, namely, they vanish for the Riemannian metrics.

Take an arbitrary plane $P \subset T_x M$ (flag) and a non-zero vector $y \in P$ (flag pole), the *flag curvature* K(P, y) is defined by

$$\mathbf{K}(P,y) := \frac{g_y(\mathbf{R}_y(v),v)}{g_y(y,y)g_y(v,v) - g_y(v,y)g_y(v,y)}.$$

We say that a Finsler metric F is of scalar curvature if for any $y \in T_x M$, the flag curvature $\mathbf{K} = \mathbf{K}(x, y)$ is a scalar function on TM_0 . If $\mathbf{K} = constant$, then F is said to be of constant flag curvature. The important of the quantity \mathbf{H} lies in the following well-known theorem:

Theorem 2.1. ([1]) Let F be a Finsler metric of scalar flag curvature on an n-dimensional manifold M (n > 2). Then the flag curvature $\mathbf{K} = constant$ is a scalar function on M if and only if $\mathbf{H} = 0$.

Let (M, F) be an n-dimensional Finsler space. For every $x \in M$, assume that

$$S_x M = \Big\{ y \in T_x M | F(x, y) = 1 \Big\}.$$

 S_xM is called the indicatrix of F at $x \in M$ and it is a compact hyper surface of T_xM , for every $x \in M$. Let $v: S_xM \hookrightarrow T_xM$ be its canonical embedding, where ||v|| = 1. Let (t, U) be a coordinate system on S_xM . Then, S_xM is represented locally by $v^i = v^i(t^{\eta}), \ \eta = 1, 2, ..., (n-1)$. One can show that

$$\frac{\partial}{\partial v^i} = F \frac{\partial}{\partial y^i}.$$

The (n-1) vectors $\{(v_{\eta}^i)\}$ form a basis for the tangent space of $S_x M$ in each point, where

$$v^{i}{}_{\eta}:=\frac{\partial v^{i}}{\partial t^{\eta}}, \quad \eta=1,2,...,(n-1).$$

For the sake of simplicity, put

$$\partial_\eta := \frac{\partial}{\partial t^\eta}$$

and observe that

$$\partial_{\eta} = F v^{i}{}_{\eta} \frac{\partial}{\partial y^{i}}.$$

Let $g = g_{ij}(x, y)dy^i dy^j$ is a Riemannian metric on $T_x M$. Inducing g on $S_x M$, one gets the Riemannian metric

$$\bar{g} = \bar{g}_{\eta\gamma} dt^{\eta} dt^{\gamma},$$

where

$$\bar{g}_{\eta\gamma} := v^i{}_{\eta}v^i{}_{\gamma}g_{ij}$$

The canonical unit vertical vector field $V(x, y) = \ell^i \frac{\partial}{\partial y^i}$ together the (n-1) vectors ∂_η , form the local basis for $T_x M$, $\mathcal{B} = \{u^1, u^2, ..., u^n\}$, where, $u^\eta = (v_{\eta}^i)$ and $u^n = V$. We conclude that

$$g(V,\partial_{\eta})=0,$$

that is to say that

$$y_i v^i{}_\eta = 0.$$

Theorem 2.2. If $F = \frac{\alpha^2}{\alpha - \beta}$ be an Einstein metric, then the following statements hold:

(a) F is Ricci-flat.

(b) α is Ricci-flat.

(c) β is constant Killing and $s^k_{0|k} = 0$.

Consider the following conventions in notations:

$$\begin{aligned} q_{ij} &:= r_{im} s_j^m, \\ t_{ij} &:= s_{im} s_j^m, \\ t_j &:= b^i t_{ij} = s_m s_j^m, \\ A_k &:= 2cs_k + c^2 b_k + t_k + \frac{1}{2}c_k \\ \Psi_k &:= 3c^2 y_k - c^2 \beta b_k + 2\beta c_k - c_0 b_k + s_0 s_k + 2s_{0|k} - s_{k|0} - 6cs_{k0}, \end{aligned}$$

where, c = c(x) is a scalar function and $c_k = \partial c / \partial x^k$. Notice that

$$y_k := a_{jk} y^j$$
 and $y_0 = \alpha^2$.

In [9], Li and Shen proved the following characterization of the Ricci-quadratic Randers metrics.

Theorem 2.3. [9] Let $F = \alpha + \beta$ be a Randers metric on an n-manifold. Then it is Ricci-quadratic if and only if

$$r_{00} = c(\alpha^2 - \beta^2), \tag{2.1}$$

$$s^{k}_{\ 0|k} = A_0,$$
 (2.2)

where, c = c(x) is a scalar function. In this case,

$$\operatorname{\mathbf{Ric}} = \overline{\operatorname{\mathbf{Ric}}} - 2t_{00} - t^{k}_{\ k}\alpha^{2} + (n-1)\Psi_{0}.$$

$$(2.3)$$

3. Proof of Theorems

Now, we ready to prove Theorem 1.1.

Proof of Theorem 1.1: Denote the Riemann curvature of α by \tilde{R}^i_{jkl} . Using the Ricci identity for $\operatorname{\mathbf{Ric}}_{ij}$, with respect to $\tilde{\nabla}$, one obtains

$$\operatorname{\mathbf{Ric}}_{ij|l|k} - \operatorname{\mathbf{Ric}}_{ij|k|l} = -\operatorname{\mathbf{Ric}}_{rj} \tilde{R}^{r}_{\ ikl} - \operatorname{\mathbf{Ric}}_{ir} \tilde{R}^{r}_{\ jkl} - \frac{\partial \operatorname{\mathbf{Ric}}_{ij}}{\partial y^{r}} \tilde{R}^{r}_{\ 0kl}.$$
 (3.1)

Multiply (3.1) by y^i , we get

$$\operatorname{\mathbf{Ric}}_{0j|l|k} - \operatorname{\mathbf{Ric}}_{0j|k|l} = -\operatorname{\mathbf{Ric}}_{rj} \tilde{R}^{r}_{0kl} - \operatorname{\mathbf{Ric}}_{0r} \tilde{R}^{r}_{jkl}.$$
(3.2)

One can easily observe that

$$\operatorname{Ric}_{ij|0} = \operatorname{Ric}_{0j|i} = \operatorname{Ric}_{i0|j} = 0.$$
(3.3)

Multiplying (3.2) by y^l and using (3.3) we obtain

$$\mathbf{Ric}_{0j|0|k} - \mathbf{Ric}_{0j|k|0} = -\mathbf{Ric}_{rj} \ \tilde{R}^{r}_{\ 0k0} - \mathbf{Ric}_{0r} \ \tilde{R}^{r}_{\ jk0} = 0.$$
(3.4)

It results immediately that

$$\frac{1}{2} \frac{\partial^2 \mathbf{Ric}}{\partial y^r \partial y^j} \ \tilde{R}^r_{\ 0k0} + \frac{\partial \mathbf{Ric}}{\partial y^r} \ \tilde{R}^r_{\ jk0} = 0.$$
(3.5)

Multiplying (3.5) by a^{jk} yields

$$\frac{1}{2} \frac{\partial^2 \mathbf{Ric}}{\partial y^r \partial y^k} \tilde{R}^r{}_0{}^k{}_0 + a^{jk} \tilde{R}^r{}_{jk0} \frac{\partial \mathbf{Ric}}{\partial y^r} = 0$$
(3.6)

Define the operator Υ as follows

$$\Upsilon := \tilde{R}^{r}{}_{0}{}^{k}{}_{0} \frac{1}{2} \frac{\partial^{2}}{\partial y^{r} \partial y^{k}} + a^{jk} \tilde{R}^{r}{}_{jk0} \frac{\partial}{\partial y^{r}}.$$
(3.7)

Let us put

$$\rho := \alpha^{-2} \mathbf{Ric}$$

Then we have

$$\partial_{\eta}\rho = \alpha^2 v^i{}_{\eta}\rho_{;i}, \qquad (3.8)$$

and

$$\partial_{\beta}\partial_{\eta}\rho = \alpha \partial_{\beta}v^{i}{}_{\eta} \rho_{;i} + \alpha^{2}v^{i}{}_{\eta}v^{j}{}_{\beta} \rho_{;i;j} + \alpha v^{j}{}_{\beta}(v^{i}{}_{\eta;j})\rho_{;i}.$$
(3.9)

Since

$$v^{j}_{\ \beta} \ \frac{\partial \alpha}{\partial y^{j}} = 0,$$

then we get

$$\partial_{\beta}\partial_{\eta}\rho = \alpha \partial_{\beta} v^{i}_{\ \eta} \ \rho_{;i} + \alpha^{2} v^{i}_{\ \eta} v^{j}_{\ \beta} \ \rho_{;i;j}.$$
(3.10)

Multiplying the two sides of (3.10) by

$$\tilde{R}^{\alpha\beta} := \tilde{R}^{\alpha}{}^{\beta}{}_{n}{}^{n}{}_{n}$$

we obtain

$$\tilde{R}^{\eta\beta}\partial_{\beta}\partial_{\eta}\rho = \tilde{R}^{i\ j}{}_{0\ 0}{}_{0}\rho_{;i;j} + \alpha \tilde{R}^{\eta\beta}\partial_{\beta}v^{i}{}_{\eta}\rho_{;i}.$$
(3.11)

It follows that

$$\tilde{\Upsilon}(\rho) := \tilde{R}^{\alpha\beta} \partial_{\beta} \partial_{\alpha} \rho - B^{\alpha} \partial_{\alpha} \rho = 0, \quad (\alpha, \beta = 1, \cdots, n-1)$$
(3.12)

where

$$B^{\eta} := 2v^{\eta}_{\ i} \tilde{R}^{i}_{\ \beta n\gamma} \tilde{a}^{\beta\gamma} - \alpha \tilde{R}^{\beta\gamma} \ \partial_{\gamma} v^{\eta}_{\ \beta}.$$

Assuming the equation (3.12) on each indicatrix $S_x M$ and using the maximum principle of Hopf, we find ρ as a function of x, only. Therefore, there is a function c(x) such that

$$\mathbf{Ric} = c(x)\alpha^2.$$

Since it must satisfy $\operatorname{\mathbf{Ric}}_{|_0} = 0$, it results that, the function c(x) is a constant and the relation

$$\mathbf{Ric} = c\alpha^2$$

holds for some constant $c \in \mathbb{R}$. The converse is also true, since by a simple calculation we have $\operatorname{\mathbf{Ric}}_{ij|0} = 0$.

By the Theorem 1.1, we obtain a necessary and sufficient condition for an Einstein (α, β) -metric to be a Riemannian metric.

Corollary 3.1. Let F be an Einstein metric on a connected semi-Riemannian manifold (M, α) . Suppose that α is of positive (negative) sectional curvature and $\operatorname{Ric}(x, y) \neq 0$. Then F is Riemannian if and only if $\operatorname{Ric}_{ij|_0} = 0$.

Proof. By Theorem 1.1, we have

$$\mathbf{Ric} = c\alpha^2,$$

where $c \in \mathbb{R}$. Since F is an Einstein metric, we have

$$\mathbf{Ric} = (n-1)\sigma F^2,$$

where $\sigma = \sigma(x)$ is a function on M. Therefore F is conformal to the Riemannian metric α , i.e, F is a Riemannian metric. The converse is trivial.

Remark 3.2. The family of Randers metrics on S^3 constructed by Bao-Shen are weakly Berwald which are not Berwaldian [6][19]. Denote generic tangent vectors on S^3 as

$$u\frac{\partial}{\partial x} + v\frac{\partial}{\partial y} + w\frac{\partial}{\partial z}$$

The Finsler function for Bao-Shen's Randers space is given by

$$F(x, y, z; u, v, w) = \alpha(x, y, z; u, v, w) + \beta(x, y, z; u, v, w)$$

with

$$\alpha = \frac{\sqrt{K(cu - zv + yw)^2 + (zu + cv - xw)^2 + (-yu + xv + cw)^2}}{1 + x^2 + y^2 + z^2}$$

$$\beta = \frac{\pm\sqrt{K-1} (cu - zv + yw)}{1 + x^2 + y^2 + z^2},$$

where K > 1 is a real constant. This family of Randers metrics are Einstein metrics of positive sectional curvature and have $\operatorname{Ric}_{ij|_0} \neq 0$, while they are not Riemannian manifolds.

Proof of Theorem 1.2: The curvature form of Berwald connection is given by

$$\Omega^{i}{}_{j} = d\omega^{i}{}_{j} - \omega^{k}{}_{j} \wedge \omega^{i}{}_{k} = \frac{1}{2}R^{i}{}_{jkl}\omega^{k} \wedge \omega^{l} - B^{i}{}_{jkl}\omega^{k} \wedge \omega^{n+l}.$$
(3.13)

For the Berwald connection, we have the following structure equation:

$$dg_{ij} - g_{jk}\Omega^k_{\ i} - g_{ik}\Omega^k_{\ j} = -2L_{ijk}\omega^k + 2C_{ijk}\omega^{n+k}.$$
(3.14)

Differentiating (3.14) yields the following Ricci identity:

$$g_{pj}\Omega^{p}_{\ i} - g_{pi}\Omega^{p}_{\ j} = -2L_{ijk|l}\omega^{k}\wedge\omega^{l} - 2L_{ijk,l}\omega^{k}\wedge\omega^{n+l} - 2C_{ijl|k}\omega^{k}\wedge\omega^{n+l} - 2C_{ijl|k}\omega^{k}\wedge\omega^{n+l} - 2C_{ijl}\Omega^{p}_{\ l}y^{l}.$$
(3.15)

It follows from (3.15) that:

$$C_{ijl|k} + L_{ijk,l} = \frac{1}{2}g_{pj}B^{p}_{\ ikl} + \frac{1}{2}g_{ip}B^{p}_{\ jkl}.$$
(3.16)

Differentiating of (3.13) yields:

$$d\Omega_i^{\ j} - \omega_i^{\ k} \wedge \Omega_k^{\ j} + \omega_k^{\ j} \wedge \Omega_i^{\ k} = 0.$$
(3.17)

Define $B^{i}_{\ jkl|m}$ and $B^{i}_{\ jkl,m}$ by:

$$dB^{i}_{jkl} - B^{i}_{mkl}\omega^{m}_{i} - B^{i}_{jml}\omega^{m}_{k} - B^{i}_{jkm}\omega^{m}_{l} + B^{i}_{jkl}\omega^{i}_{m} = B^{i}_{jkl|m}\omega^{m} + B^{i}_{jkl,m}\omega^{n+m}.$$
(3.18)

Similarly, we define $R^i_{\ jkl|m}$ and $R^i_{\ jkl,m}$:

$$dR^{i}_{jkl} - R^{i}_{mkl}\omega^{m}_{i} - B^{i}_{jml}\omega^{m}_{k} - R^{i}_{jkm}\omega^{m}_{l} + R^{i}_{jkl}\omega^{i}_{m} = R^{i}_{jkl|m}\omega^{m} + R^{i}_{jkl,m}\omega^{n+m}.$$
(3.19)

From (3.17), (3.18) and (3.19), one obtain the following Bianchi identities:

$$R^{i}_{jkl|m} + R^{i}_{jlm|k} + R^{i}_{jmk|l} = 0, ag{3.20}$$

$$B^{i}_{\ jkl|m} - B^{i}_{\ jkm|l} = R^{i}_{\ jkl,m}, \tag{3.21}$$

$$B^i_{jkl,m} = B^i_{jkm,l}. (3.22)$$

Contracting i and k in (3.21) yields

$$B^{p}_{\ jpl|m} - B^{p}_{\ jpm|l} = R^{p}_{\ jpl,m}.$$
 (3.23)

By definition of the Riemann curvature of Berwald connection, we have

$$R^{i}_{\ jkl}(x,y) = \frac{1}{3} \frac{\partial}{\partial y^{j}} \Big\{ \frac{\partial R^{i}_{k}}{\partial y^{l}} - \frac{\partial R^{i}_{l}}{\partial y^{k}} \Big\}.$$
(3.24)

Following (3.24) a Finsler space is of quadratic Riemann curvature if and only of the Berwald-Riemann curvature depends only to the position x. Now we have

$$R^{i}_{\ k} = R^{i}_{\ jkl}(x,y)y^{j}y^{l},$$

$$\mathbf{Ric} = R^{p}_{\ jpl}(x,y)y^{j}y^{l}.$$
(3.25)

We get

Then **Ric** is quadratic in $y \in T_x M$ if R^p_{jpl} are functions of position alone, i.e., $R^p_{jpl} = R^p_{jpl}(x)$. This yields

$$R^{p}_{\ jpl,m} = 0. (3.26)$$

By (3.23) and (3.26) we have

$$B^p_{jpl|m} = B^p_{jpm|l}. (3.27)$$

Multiplying (3.27) with y^m

$$E_{jk|m}y^m = 0. (3.28)$$

This completes the proof.

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Azadeh Shirafkan

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